Non-Linear Thermal Conductance and Noise Dependence INCT- SC - 23/04/2013

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References:

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PRE 79 051116 (2009)
JSTAT P06010 (2011)
Physica A 391 3816 (2012)
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Abstract

- Simple small models can be quite illustrative about the way biological and molecular systems behave.
- We show how the couplings between the model's non-linearities and non-equilibrium properties of thermal noise lead to new properties of transport coefficients.

Why are we interested in this problem?

- Under certain conditions, there might be non-equilibrium energy feeding into a system (mostly small), i.e, non-gaussian noises will act upon these systems. For instance, Poisson noise.
- We might be able to solve small system models with arbitrary noise properties.
- What happens when non-linearities couple with the non-gaussian (higher cumulants) properties of noise?

Equilibrium noise

- An equilibrium thermal reservoir interacts with a system, say a Brownian particle, by an effective interaction that may be well represented by white gaussian noise.
- That form of noise ensures the Boltzmann-Gibbs equilibrium distribution for the system.
- Only the mean and variance of the noise exists.
- For non-equilibrium noise, higher order cumulants may exist.

Non-Equilibrium Stationary State - NESS

- A system in out-of equilibrium conditions might reach a NESS.
- This can be done by submitting the system to contact with distinct temperature reservoirs, or by submitting the system to a non-equilibrium noise, e.g. Poisson noise.
- We shall see that the distinction above is less than it seems.

Poisson noise & Brownian particles

A Brownian particle under Poisson noise is a paradigmatic NESS for $t \to \infty$.

The model is

$$\dot{x}(t)=v(t),$$

$$M\dot{v}(t) = -k_0x(t) - \gamma v(t) + \eta(t),$$

where η is a Poisson noise,

$$\eta(t) = \sum_{\ell} \Phi(t) \delta(t - t_{\ell}).$$

It is interesting to chose a variable rate

$$\lambda(t) = \lambda_0 \left[1 + A \cos(\omega t) \right],$$



Noise Cumulants

Poisson noise has infinite number of cumulants:

$$\langle \eta_{i_1}(t_1) \eta_{i_2}(t_2) \dots \eta_{i_n}(t_n) \rangle_c = \overline{\Phi^n} \lambda(t_1) \delta_{i_1 i_2} \dots \delta_{i_{n-1} i_n}$$

$$\times \delta(t_1 - t_2) \dots \delta(t_{n-1} - t_n).$$

The general Laplace transform for noise reads:

$$\langle \tilde{\eta}_1(s_1) \dots \tilde{\eta}_n \rangle_c = \frac{A_n}{s_1 + \dots + s_n},$$

where A_n is the n-th order cumulant

• Energy injection, dissipation, and the received heat $(x_0 = v_0 = 0)$:

$$Q = \int_{0}^{\tau} \left[\eta(t) \ v(t) - \gamma v(t)^{2} \right] dt = \frac{1}{2} M \ v(t)^{2} \Big|_{0}^{\tau} + \frac{1}{2} k_{0} \ x(t)^{2} \Big|_{0}^{\tau}.$$



The time averaged exact NESS energy:

Then finally, we write the contributions for the injection of energy as

$$J_{\mathrm{ITO}}(au) = \left(rac{\lambda_0\,ar{\Phi}^2}{M} + rac{A^2\omega^2\, heta\,\lambda_0^2\,ar{\Phi}^2}{M\,\left[\omega^2\left(\omega^2 - 2\,\omega_0^2 + 4\, heta^2
ight) + \omega_0^4
ight]}
ight)\, au,$$

and

$$J_{\text{IT,cosc}}(\tau) = \frac{\lambda_0 \bar{\Phi}^2 A \sin(\omega \tau)}{M \omega} + \frac{A \left[A \left(\omega_0^2 - \omega^2 \right) \cos \left(2 \omega \tau \right) + 4 \omega_0^2 \cos(\omega \tau) \right] \lambda_0^2 \bar{\Phi}^2}{4 M \left[\omega^2 \left(\omega^2 - 2 \omega_0^2 + 4 \theta^2 \right) + \omega_0^4 \right]} + \frac{A \left[2 \omega \theta \left(A \sin \left(2 \omega \tau \right) + 4 \sin \left(\omega \tau \right) \right) - 4 \omega^2 \cos(\omega \tau) \right] \lambda_0^2 \bar{\Phi}^2}{4 M \left[\omega^2 \left(\omega^2 - 2 \omega_0^2 + 4 \theta^2 \right) + \omega_0^4 \right]}, \quad (47)$$

and

$$J_{\text{IT}c} = \frac{\lambda_0^2 \bar{\Phi}^2}{M \omega_0^2} + \frac{A \lambda_0^2 \bar{\Phi}^2 (\omega_0^2 - \omega^2)}{M [\omega^2 (\omega^2 - 2 \omega_0^2 + 4 \theta^2) + \omega_0^4]} + \frac{[\omega^4 (\omega^2 - 12 \theta^2) + \omega_0^2 (3 \omega_0^4 - 5 \omega_0^2 \omega^2 + \omega^4 - 4 \theta^2)] A^2 \lambda_0^2 \bar{\Phi}^2}{4 M [\omega^2 (\omega^2 - 2 \omega_0^2 + 4 \theta^2) + \omega_0^4]^2}.$$
 (48)

For the total energy flux $J_E = J_{\rm IT} + J_{\rm DT}$ it can be easily seen that the linear term on τ cancels out, while the constant term becomes exactly the average energy of equation (40). The non-oscillating part of the energy is

$$J_{\rm E}^{\rm no} = \frac{\bar{\Phi}^2 \lambda_0 (2 M \omega_0^2 + \lambda_0 \gamma)}{2 \gamma M \omega_0^2} + \frac{(\omega^2 + \omega_0^2) A^2 \bar{\Phi}^2 \lambda_0^2}{4 M \left[\omega^2 (\omega^2 - 2 \omega_0^2 + 4 \theta^2) + \omega_0^4\right]},$$
 (49)

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Ressonating Energy Injection 2

Injected - Dissipated Energy = Heat

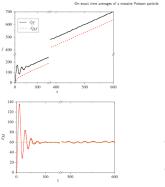
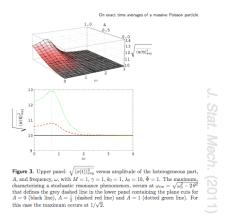


Figure 4. Upper panel: total injected and (symmetric) dissipated power, $J_{\rm el}(-J_{\rm el})$ versus time, $\tau_{\rm e}$ exceeding to the definition in the largend. Lower panel: total energy, $E_{\rm el}$, versus time, $\tau_{\rm e}$. The dashed (green) line represents the anymptotic limit plum by equation (oil), in both cases we have used the following values: M = 10, $k_0 = 1$, $\gamma_{\rm el} = 1$, $\beta_{\rm e$

Ressonating Energy Injection 3

The oscillation amplitude is modulated by the resonance between the natural frequency $\omega_0 = \sqrt{\frac{k}{m}}$ and the poisson frequency ω .



Stationary Distributions



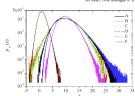


Figure 1. Numerically obtained probability density function $p_0(x)$ versus position x for various cases with $\lambda = 10$, $\delta = 1$ and the noise defined by equation (d) with $\omega = x$. Following the legend in the figure we have the respective cases, λ : $M = 1, k_0 = 1, \gamma = 1, A = 0$, B: $M = 10, k_0 = 1, \gamma = 1, A = 0$, C: $M = 10, k_0 = 1, \gamma = 1, A = 0$, E: $M = 10, k_0 = 1, A_0 = 1, A_0$

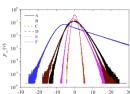


Figure 2. Numerically obtained probability density function $p_{ss}(v)$ versus scalar velocity v for the same parameter sets of figure 1.

J. Stat. Mech. (2011) P06010

Linear Poisson model and energy

- We did not explore fully the Poisson model due to its linearity.
- We could not couple the energy with the higher order Poisson cumulants.
- The stationary state is a NESS one: the probability distribution is not Boltzmann-Gibbs, hence NE.

Time average

• We start from (1D Brownian Particle)

$$p(x, v, t) = \langle \delta(x - x(t)) \, \delta(v - v(t)) \rangle,$$

where x(t) and v(t) are solutions for the equations of motion given a realization of the noise.

Time averaging:

$$p^{ss}(x,v) = \lim_{s \to 0} s \int_0^\infty dt \, e^{-st} \, \left\langle \delta(x-x(t)) \, \delta(v-v(t)) \right\rangle.$$

After a straightforward calculation...

$$p^{ss}(x,v) = \lim_{s \to \infty} \lim_{\epsilon \to \infty} \int_{-\infty}^{\infty} \frac{dQ}{2\pi} e^{iQx} \int_{-\infty}^{\infty} \frac{dP}{2\pi} e^{iPv} \times$$

$$\times \sum_{n=0}^{\infty} \sum_{l=0}^{\infty} \frac{(-iQ)^n}{n!} \frac{(-iP)^l}{l!} \times$$

$$\times \int_{-\infty}^{\infty} \frac{dq_1}{2\pi} \dots \int_{-\infty}^{\infty} \frac{dq_n}{2\pi} \int_{-\infty}^{\infty} \frac{dp_1}{2\pi} \dots \int_{-\infty}^{\infty} \frac{dp_l}{2\pi} \times$$

$$\times \frac{s}{s - (iq_1 + \epsilon + \dots + iq_n + \epsilon + ip_1 + \epsilon + \dots + ip_l + \epsilon)} \times$$

$$\times \langle \tilde{x}(iq_1 + \epsilon) \dots \tilde{x}(iq_n + \epsilon) \tilde{v}(ip_1 + \epsilon) \dots \tilde{v}(ip_l + \epsilon) \rangle.$$

Cumulant-energy coupling

• Due to the linearity of the model $(\tilde{v}(s) = s \, \tilde{x})$

$$x, v(s) = \frac{\tilde{\eta}(s)}{R(s)}.$$

• The exchanged heat Q is quadratic in $\tilde{\eta}$, hence

$$Q \propto \langle \tilde{\eta}(s_1) \, \tilde{\eta}(s_2) \rangle$$
.

 We see that heat can only be a function of the variance, and average, of the noise

$$\langle \tilde{\eta}(s_1) \, \tilde{\eta}(s_2) \rangle = \langle \tilde{\eta}(s_1) \, \tilde{\eta}(s_2) \rangle_c + \langle \tilde{\eta}(s_1) \rangle_c \, \langle \tilde{\eta}(s_2) \rangle_c$$



Non-linear Interactions

• For a model with non-linear springs (non-linear force $= -k_3 x^3$) we can write:

$$R(s)\tilde{x}(s)+k_3\int_{-\infty}^{\infty}\frac{dq_1}{2\pi}\frac{dq_1}{2\pi}\frac{dq_1}{2\pi}\frac{\tilde{x}(iq_1)\tilde{x}(iq_2)\tilde{x}(iq_3)}{s-i(q_1+q_2+q_3)}=\tilde{\eta}(s).$$

- The recurrence relation above shows that we can now couple the non-linearity with the higher order cumulants.
- In first order on k₃:

$$\langle \tilde{x}(s_1) \, \tilde{x}(s_2) \rangle \rightarrow k_3 \int_{-\infty}^{\infty} \frac{dq_1}{2\pi} \, \frac{dq_1}{2\pi} \, \frac{dq_1}{2\pi} \, \frac{\langle \tilde{x}(s_1) \, \tilde{x}(iq_1) \tilde{x}(iq_2) \tilde{x}(iq_3) \rangle}{s_1 - i(q_1 + q_2 + q_3)}$$

where we can express $\langle \tilde{x}(s_1) \tilde{x}(iq_1) \tilde{x}(iq_2) \tilde{x}(iq_3) \rangle$ as noisy combinations of averages, variances, skewness and kurtosis!



Higher order cumulants as distinct reservoirs

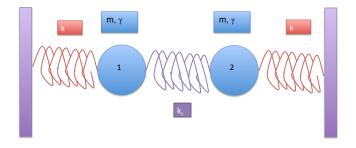
• Assuming the odd cumulants to be zero, we can express:

$$\begin{split} \langle \tilde{x}(s_1) \, \tilde{x}(iq_1) \, \tilde{x}(iq_2) \, \tilde{x}(iq_3) \rangle &= \langle \tilde{x}(s_1) \, \tilde{x}(iq_1) \, \tilde{x}(iq_2) \, \tilde{x}(iq_3) \rangle_c \, + \\ &+ \langle \tilde{x}(s_1) \, \tilde{x}(iq_1) \rangle_c \, \langle \tilde{x}(iq_2) \, \tilde{x}(iq_3) \rangle_c \, + \\ &+ \langle \tilde{x}(s_1) \, \tilde{x}(iq_2) \rangle_c \, \langle \tilde{x}(iq_1) \, \tilde{x}(iq_3) \rangle_c \, + \\ &+ \langle \tilde{x}(s_1) \, \tilde{x}(iq_3) \rangle_c \, \langle \tilde{x}(iq_2) \, \tilde{x}(iq_1) \rangle_c \, . \end{split}$$

- The received heat term has the (usual) contributions from the gaussian part (variance), but also the extra (unusual) kurtosis contribution.
- Thus, variance and kurtosis act as if distinct energy reservoirs are coupled to the system.

Small systems: our model

Two connected Brownian particles; non-linear springs; two thermal reservoirs.



Langevin Equation

Equations to be solved

$$m\frac{dv_{i}(t)}{dt} = -k x_{i}(t) - \gamma v_{i}(t) - \sum_{l=1}^{2} k_{2l-1} [x_{i}(t) - x_{j}(t)]^{2l-1} + \eta_{i}(t)$$

with

$$\frac{dx_{i}(t)}{dt}=v_{i}(t)$$

where $(i,j) \in \{1,2\}$ and k_1 and k_3 are the linear and non-linear coupling constants, respectively. The system is decoupled (linear) for $k_{1,3} = 0$.

Heat flux between particles

We define the energy flux as

$$j_{12} = \frac{1}{2}(P_{21} - P_{12}) = \frac{1}{2}(F_{21} \cdot v_2 - F_{12} \cdot v_1),$$

where
$$F_{21} = -k_1(x_2 - x_1) - k_3(x_2 - x_1)^3$$
.

We have

$$j_{12}(t) \equiv -\sum_{l=1}^{2} \frac{k_{2\,l-1}}{2} \left[x_1(t) - x_2(t) \right]^{2\,l-1} \left[v_1(t) + v_2(t) \right].$$

Thermal conductance

We are finally in the position to compute the thermal conductance,

$$\kappa \equiv -\frac{\partial}{\partial \Delta T} \langle j_{12} \rangle_{\Delta T},$$

$$\kappa = -\frac{\overline{\langle j_{12} \rangle}}{T_1 - T_2}, \quad \text{(for small } T_1 - T_2\text{)}$$

Resorting to the single particle results and the equipartition theorem, we relate the cumulant features of the noise and the proper temperature, T_i , namely, $A_i(2) = 2 \gamma T_i$, yielding the thermal conductance, $\kappa = \kappa^{(0)} + \kappa^{(1)} + \mathcal{O}\left(k_3^2\right)$

Conductance: linear systems and general types of noises

 For linear systems, the higher order cumulants do not contribute. thus, the time averaged heat flux is simple and is given by

$$\overline{\left\langle j_{12}^{(B)} \right\rangle} = \overline{\left\langle j_{12}^{(0)} \right\rangle} = -\frac{k_1^2}{4} \frac{\left(\mathcal{A}_1(2) - \mathcal{A}_2(2) \right)}{m \, k_1^2 + \gamma^2 \, (k + k_1)},$$

where $A_2(2) - A_1(2) = 2 \gamma \Delta T$.

The linear contribution for the conductance is thus

$$\kappa_0 = \frac{k_1^2 \gamma}{2 (m k_1^2 + \gamma^2 (k + k_1))},$$

Conductance: non-linear systems and gaussian noises

- The reservoirs are equilibrium ones
- First order on k_3

$$\overline{\left\langle j_{12}^{(1)} \right\rangle} = -\frac{3}{8} \gamma \, k_1 \, k_3 \frac{\left(2 \, k + k_1\right) \left[\mathcal{A}_1(2)^2 - \mathcal{A}_2(2)^2 \right]}{\left(k + 2 \, k_1\right) \left[\gamma^2 \left(k + k_1\right) + m \, k_1^2 \right]^2},$$

where
$$A_1(2)^2 - A_2(2)^2 = -4 \gamma^2 \Delta T (T_1 + T_2)$$
.

 Thus, for the non-linear model and gaussian noise we have the conductance:

$$\kappa_{0,1} = \frac{k_1^2 \gamma}{2 \left(m k_1^2 + \gamma^2 (k + k_1) \right)} + \frac{3}{2} k_1 k_3 \frac{\left(2 k + k_1 \right) \gamma^3 \left(T_1 + T_2 \right)}{\left(k + 2 k_1 \right) \left[\gamma^2 (k + k_1) + m k_1^2 \right]^2}.$$

Conductance: non-linear systems and general noises

- The reservoirs are non-equilibrium ones
- The variance contribution still holds. however there is a new flux originating from the kurtosis in first order on k_3

$$\begin{split} \overline{\left\langle j_{12}^{(1,Poisson)} \right\rangle} &= -\frac{27}{2} \, \frac{\gamma^2 \, k_1 \, k_3}{\lambda} \frac{\mathcal{N}}{\mathcal{D}} \left[\mathcal{A}_1(2)^2 - \mathcal{A}_2(2)^2 \right], \\ \text{where } \mathcal{A}_1(2)^2 - \mathcal{A}_2(2)^2 &= -4 \, \gamma^2 \, \Delta \, T \, (T_1 + T_2), \text{ and} \\ \mathcal{N} &\equiv \gamma^2 \, (5 \, k + 3 \, k_1) + m \, \big(3 \, k_1^2 + 4 \, k^2 + 11 \, k \, k_1 \big) \,, \\ \mathcal{D} &\equiv \left[\gamma^2 \, (k + k_1) \right] \left[m \, (4 \, k + 9 \, k_1)^2 + 6 \, \gamma^2 \, (2 \, k + 3 \, k_1) \right] \times \\ &\qquad \times \left[3 \, \gamma^4 + m^2 \, k_1^2 + 4 \, m \, \gamma^2 \, (k + k_1) \right] \end{split}$$

Conductance: non-linear systems and general noises

Thus, for the non-linear model and general (Poisson) noise we have the conductance:

$$\kappa_{1,Poisson} = 54 \frac{\gamma^4 k_1 k_3}{\lambda} \frac{\mathcal{N}}{\mathcal{D}} (T_1 + T_2).$$

Thus, for the non-linear model and general (Poisson) noise we have the total conductance:

$$\kappa_{0,1} = \frac{k_1^2 \gamma}{2 \left(m k_1^2 + \gamma^2 (k + k_1) \right)} + \frac{3}{2} \gamma k_1 k_3 \frac{\left(2 k + k_1 \right) \gamma^2 \overline{T}}{\left(k + 2 k_1 \right) \left[\gamma^2 (k + k_1) + m k_1^2 \right] + 2}.$$

$$+54 \frac{\gamma^4 k_1 k_3}{\lambda} \frac{\mathcal{N}}{\mathcal{D}} (T_1 + T_2).$$

Observe that as the Gaussian limit is reached $(\lambda \to \infty)$, the last contribution above vanish.

System Energy

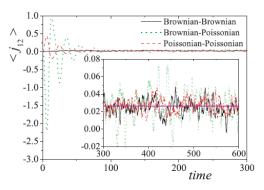


FIG. 1. (Color online) Average exchange flux $\langle j_{12} \rangle$ of a two-massive-particle system for different combinations of paradigmatic types of noise with $T_1=10$, $T_2=121/10$, m=10, $\gamma=k=1$, $k_1=1/5$, $k_3=0$, and $\lambda=10$ for Poissonian particles. After the transient, κ agrees with the theoretical value, $\kappa=21/800=0.026$ 25, with the fitting curves lying within the line thickness. The averages have been obtained by averaging over 850 \times (5 \times 10⁵) points. The discretization used is $\delta t=10^{-5}$ with snapshots at every $\Delta t=10^{-3}$.

Thermal Conductance

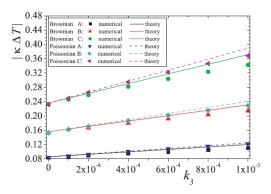


FIG. 2. (Color online) Comparison between numerically obtained values (symbols) and the first-order approximation of thermal conductance from Eqs. (8)–(10) for different temperature pairs, namely, $A = \{10, \frac{169}{10}\}$, $B = \{10, \frac{225}{10}\}$, and $C = \{10, \frac{289}{10}\}$ with m = 10, $\gamma = k = 1$, $k_1 = 1/5$, and $\lambda = 1$ for Poissonian particles.

Conclusions & perspectives

- We are working to systematize the case of a general potential.
- The method can be applied to a linear chain of arbitrary size, or a solid.
- Quantum version of the method.

Agradecimentos











